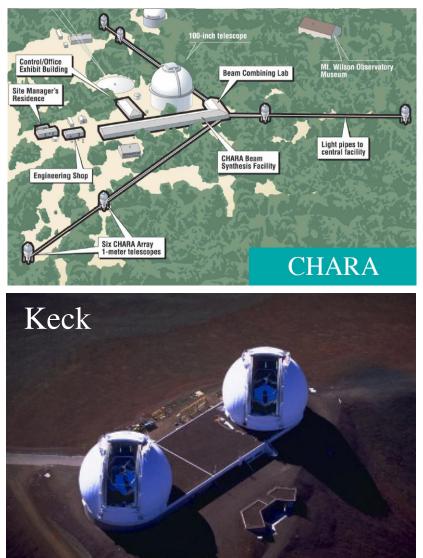


#### The Magdalena Ridge Observatory Interferometer: a high-sensitivity imaging array

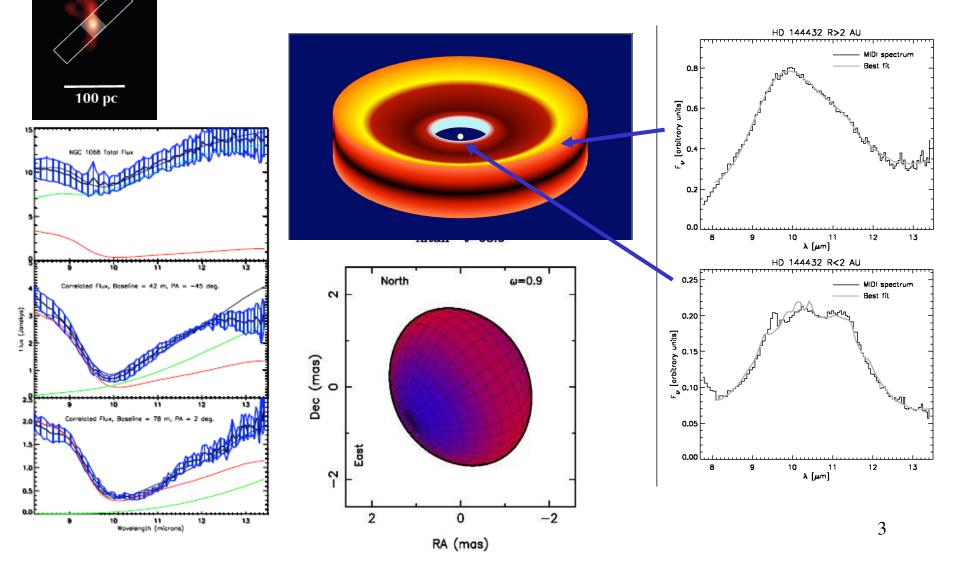
David Buscher <dfb@mrao.cam.ac.uk> and the MROI team

# The first generation of facility interferometers is now online





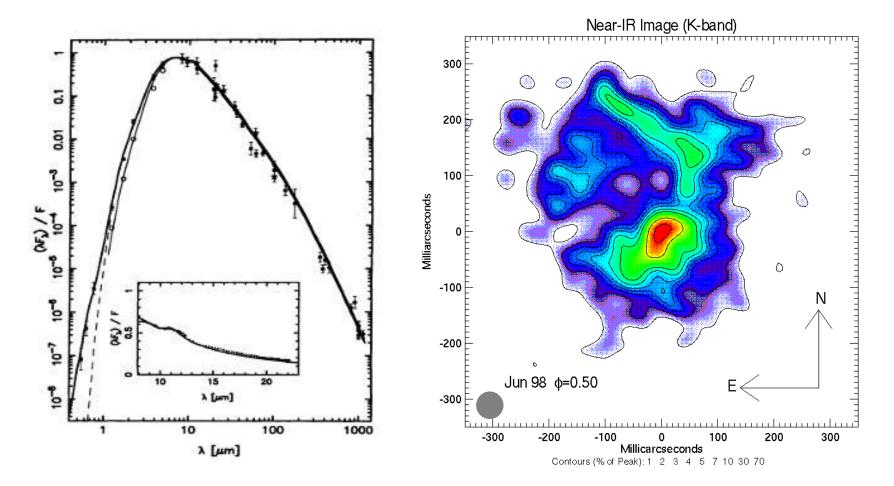
## Over 40 astrophysics papers were published in the last year alone



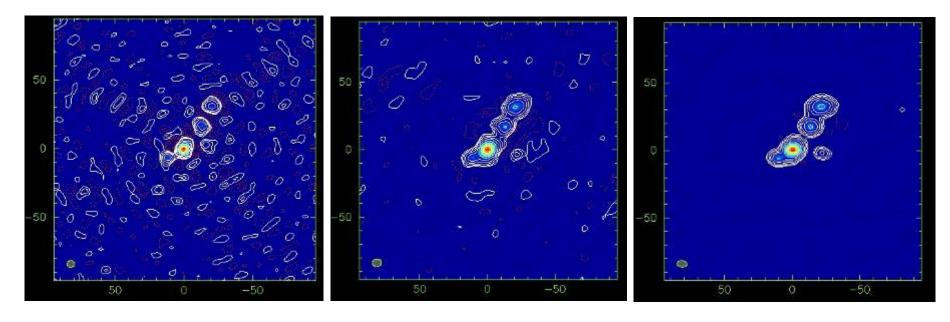
### What is needed now?

- Longer baselines?
- Bigger apertures?
- More spectral resolution?

# We need to move from modelfitting to imaging



#### Imaging needs many telescopes

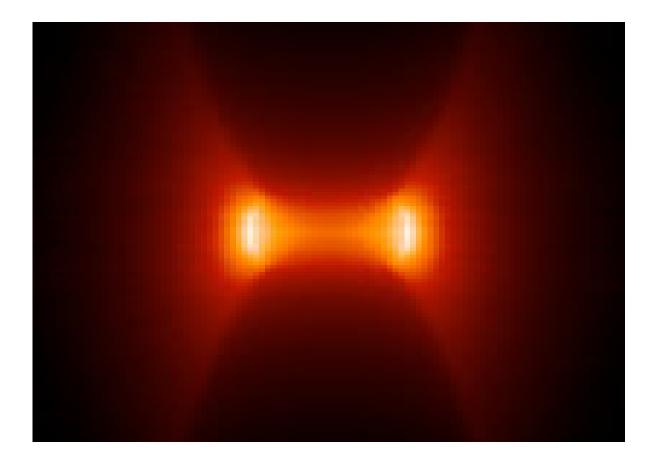


4 telescopes

6 telescopes

8 telescopes

## We need the sensitivity to access the really interesting classes of target



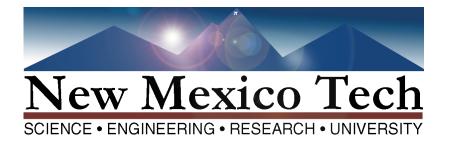
# The fringe tracker determines the faintest object we can observe

- The fringe tracker must achieve a usable signal-tonoise ratio in an integration time fixed by the atmosphere.
- If we cannot fringe-track on a target, we can never observe it interferometrically, no matter how much observing time we have.
- Using larger apertures does not help much.

#### The Magdalena Ridge Observatory is a Congressionally-funded project in New Mexico

### The MRO Interferometer is being built by a partnership between NMT and the Cavendish

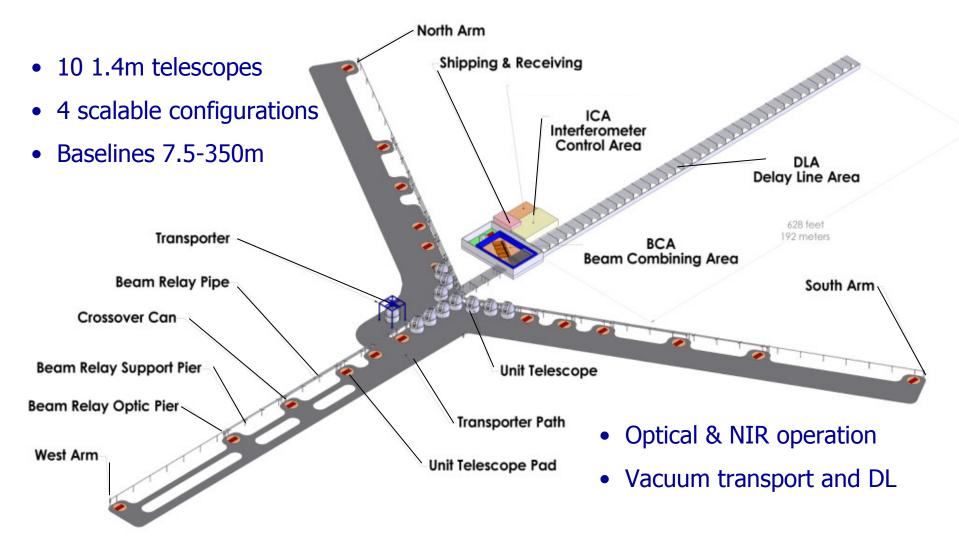




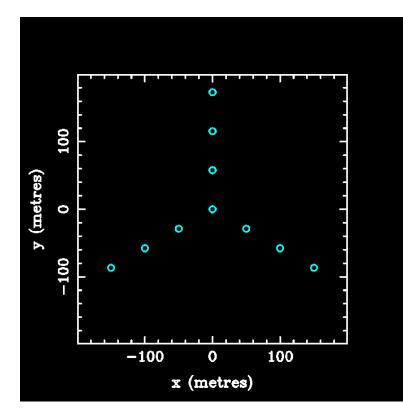
# The MROI site at 10,500ft is nearly ideal for an interferometer



# The MRO Interferometer is optimized for imaging and sensitivity



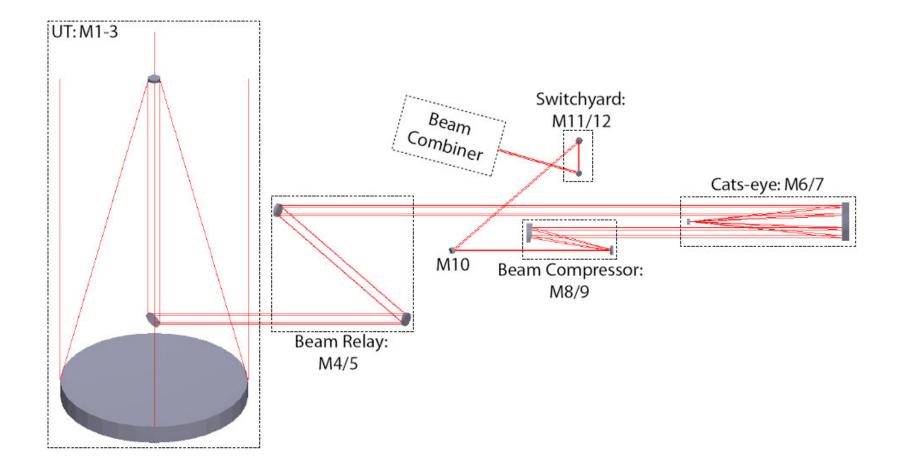
# With 10 relocatable telescopes, MROI will have unequalled imaging capability



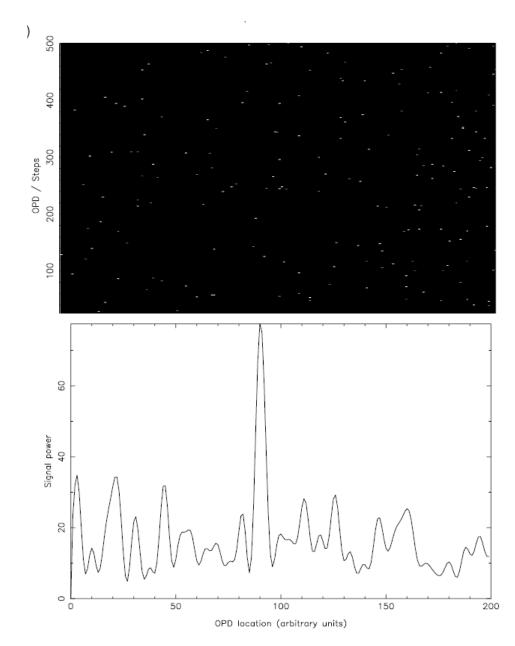
Array layout

Instantaneuous (u,v) coverage

# The beam train is optimized for maximum throughput



With group delay coherencing, it will be able to track fringes on 14<sup>th</sup> magnitude targets

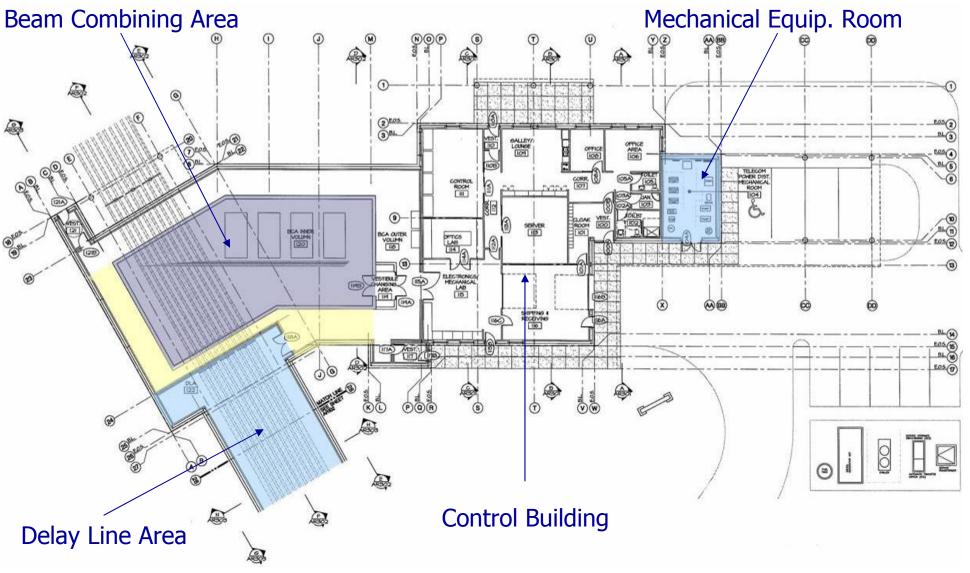


#### MROI Schedule and Team

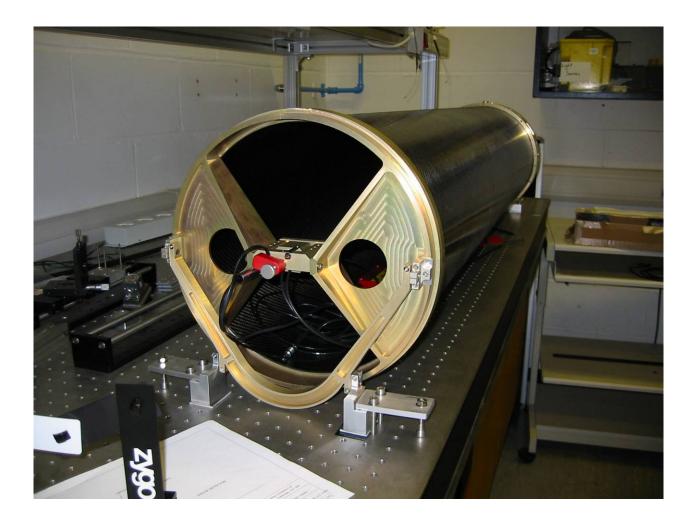
- Team
  - Manager E. Bakker
  - System Architects D. Buscher & C. Haniff
  - Scientist M. Creech-Eakman
  - Software T. Coleman
  - Controls C. Parameswariah
  - DL M. Fisher et al.
  - Beam Combiners J. Young
  - Alignment System & Optical Designs – C. Jurgenson
  - Site Characterization D.
    Klinglesmith

- Schedule
  - Buildings and infrastructure Aug '06-Aug '07
  - Telescopes Oct '07-Nov '09
  - DL Carts Oct '07
  - NIR Arrays Feb '08
  - First closure phase Aug '08
  - Commissioning thru Sept '10
  - Shared-risk science begins

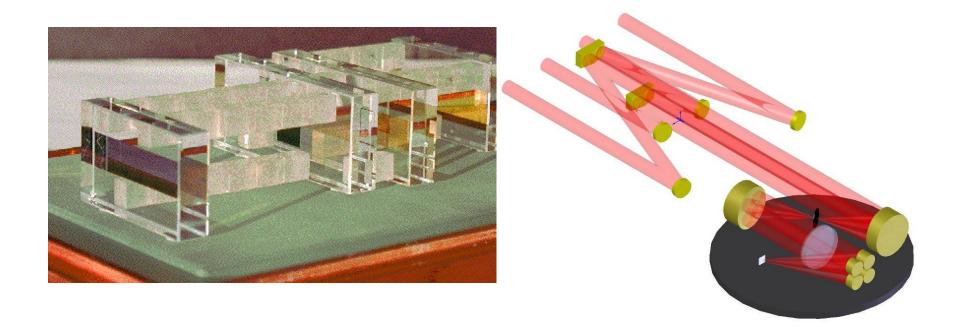
### Notice to Proceed on the interferometer buildings was issued August 4<sup>th</sup>



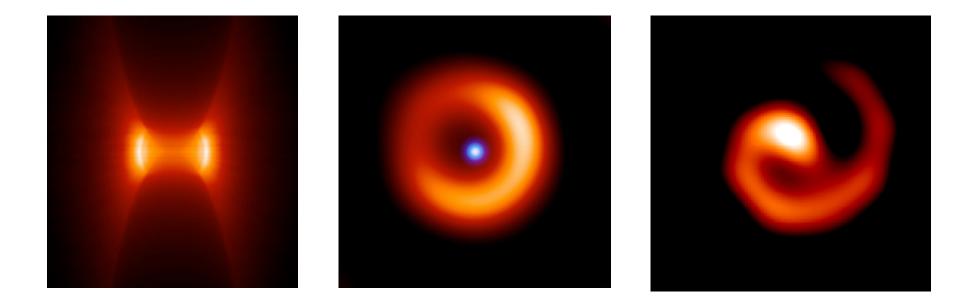
# The first delay line prototype is under construction in Cambridge



# The Beam Combiner downselect is due in November



### The MROI design is focused on 3 key Science Missions



# MROI can (in theory) obtain ~10cm spatial resolution on GEO satellites



### **Open Positions**

- Lead Opto-Mechanical Engineer
- Mechanical Engineer
- Instrument scientist
- Programmers
- Postdocs and students
- http://www.mro.nmt.edu

