

Beam Combiner studies for MROI

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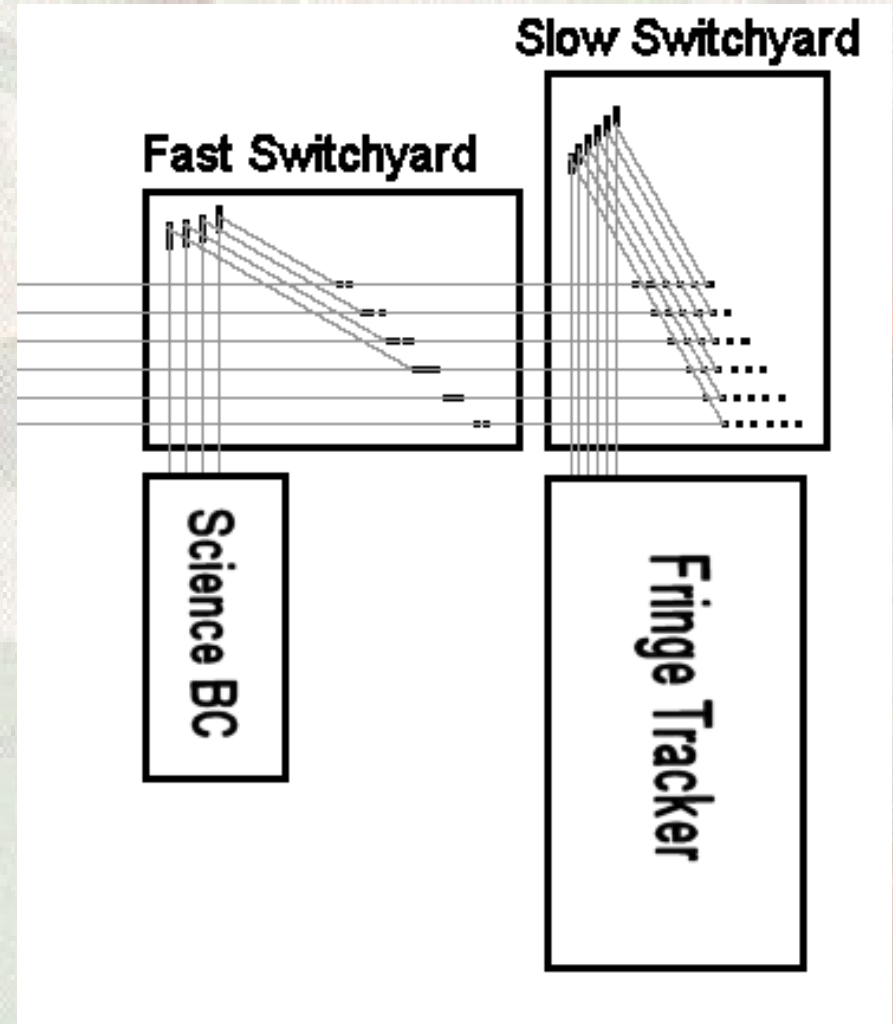
Aims & Background

- **Want beam combiners that will meet MROI top-level science goals**
- **First priority is near-infrared combiners (J,H,K bands)**
 - *Visible-wavelength combiners will be part of a later phase*
- **Cambridge currently performing studies of candidate concepts**
 - Will be finished in a few months
 - COAST experience & VLTI concept study feed into this

MROI combiner requirements

- **Top-level Requirements**

- Separate science and fringe tracking combiners
- Combiners operate simultaneously
 - Need to match paths (inter-band dispersion, internal drifts)
- Science at J or H or K, choice of $\lambda/\Delta\lambda$
- Handle 6 telescopes
- Upgrade path for 10 telescopes

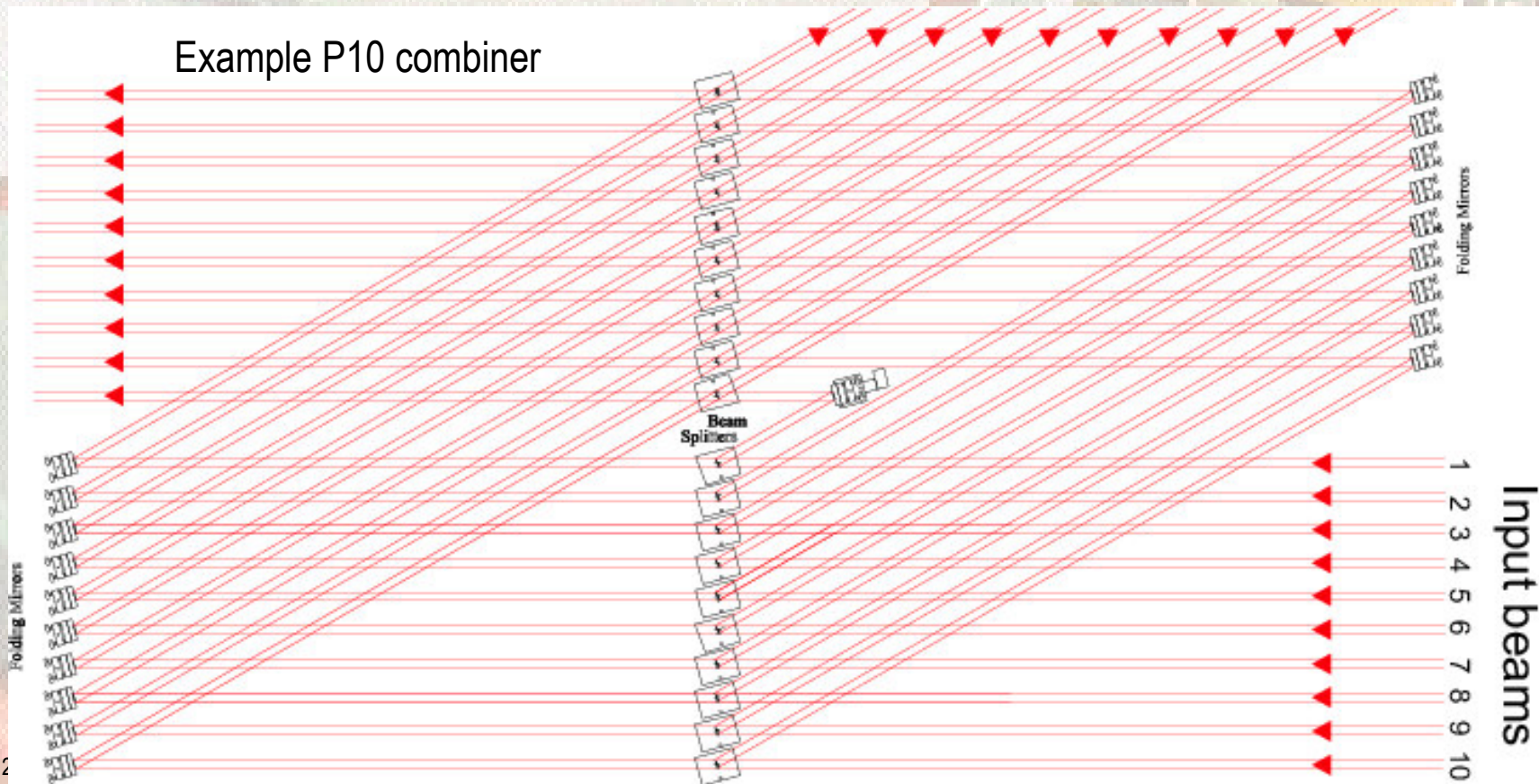


Requirements - Individual Combiners

- **Requirements for Fringe Tracking Combiner:**
 - Support hardware coherencing and software co-phasing
 - Limiting sensitivity H=14 for AGN science mission
 - Switchable between operation at H (1.5-1.8 μ m) & K (2.2-2.5 μ m)
 - **Best sensitivity will be at H**
 - Track on all nearest-neighbour baselines simultaneously
- **Top-level requirements for IR Science Combiner:**
 - Science at J (1.2-1.35 μ m) or H or K, choice of $\lambda/\Delta\lambda \sim 30, 300$ (switchable)
 - **Possibly a case for high-res. mode ($\lambda/\Delta\lambda > 1000$)**
 - S/N>2 in 100sec for K~13 at low spectral resolution
 - **For AGN science mission**
 - Access to all baselines and independent closure triangles, within ~few minutes
 - Calibration of V^2 to 2% rms, closure phase to 0.8° rms (bright targets)

Candidate Fringe Tracking Combiners

- Pupil-plane pairwise nearest neighbour (PNN)
- Image-plane pairwise nearest neighbour (INN)
- Image-plane all-on-one 6-way (I6)

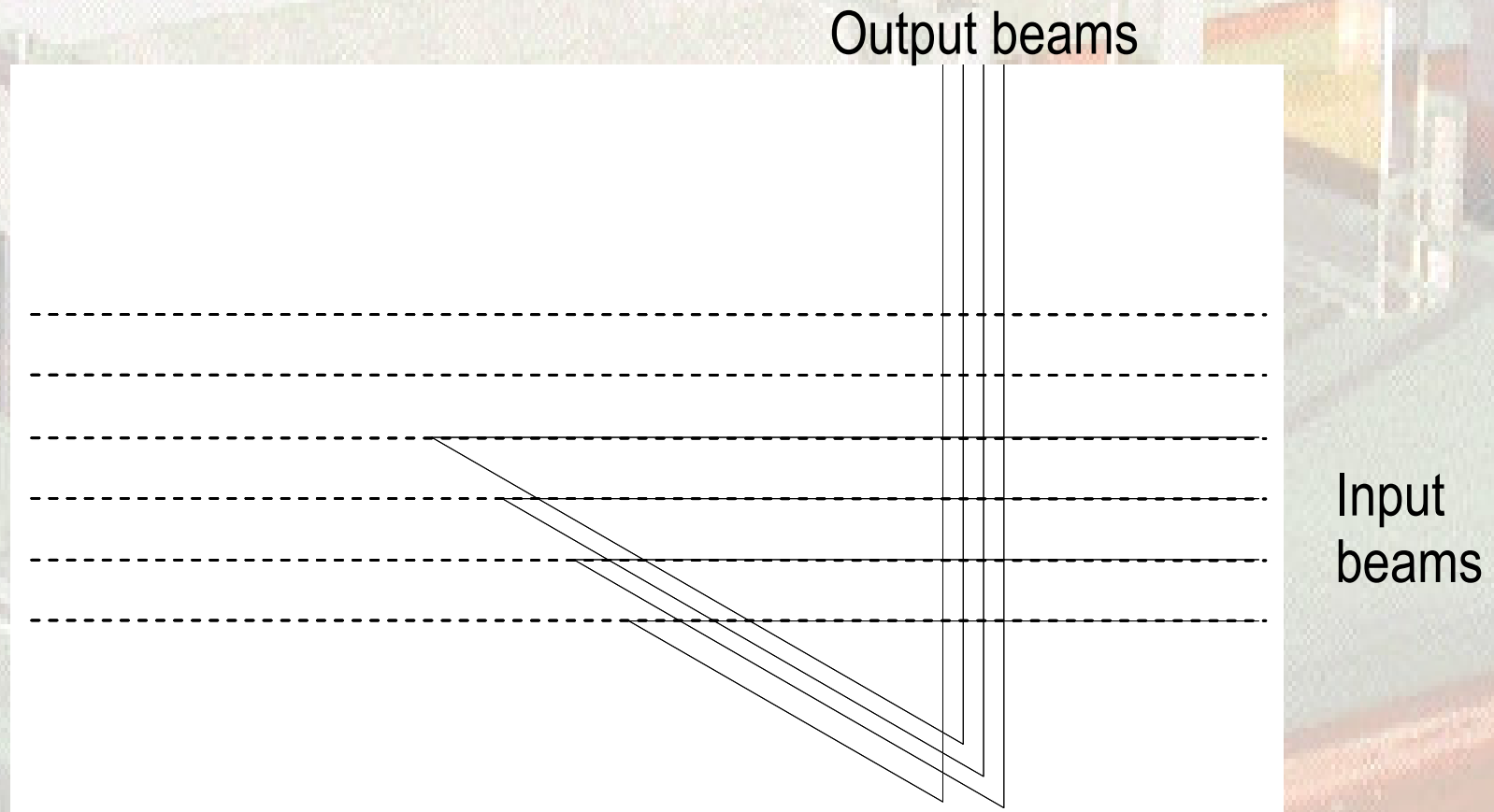


Candidate Science Combiners

- **Contacted 4-way pupil plane fed by fast switchyard (P4S)**
 - 4 of 6 beams selected by fast switchyard
 - Contacted optics combine 4 beams
 - Path modulators encode fringes in time domain
 - 4 combiner outputs, each contains all baselines
- **Contacted 8-way (6 inputs used) pupil plane (P8)**
 - Always feed in 6 beams
 - 2-position internal switchyard selects baselines/triangles measured
- **4-way image plane fed by fast switchyard (I4S)**
 - 4 of 6 beams selected by fast switchyard
 - Combine 4 beams in focal plane on single detector – spatial fringes
- **6-way image plane (I6)**

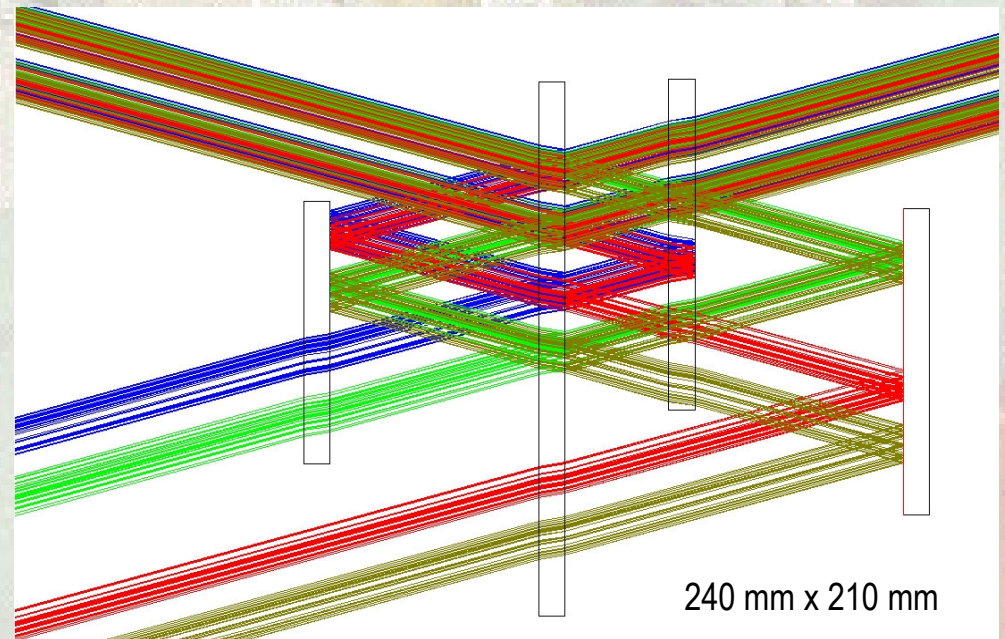
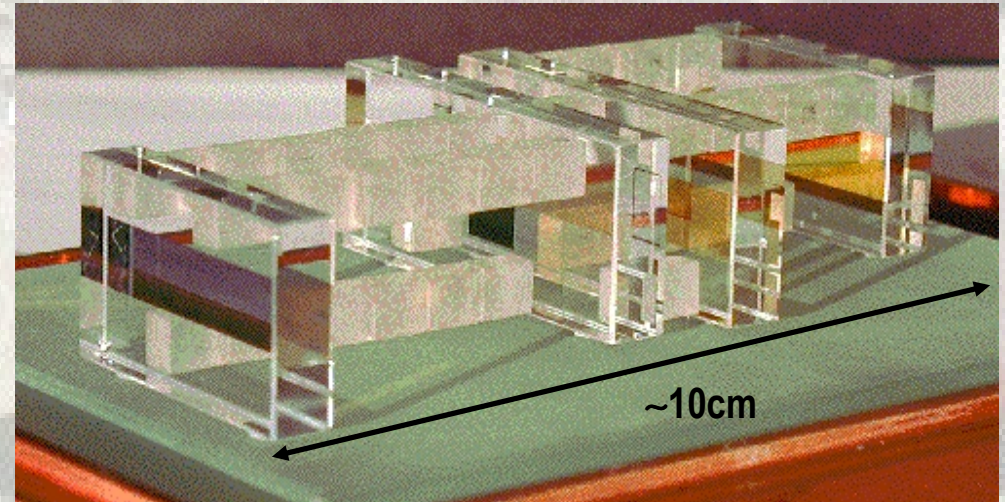
Fast Switchyard for P4S/I4S

- Needed when > 4 telescopes
- Select any 4 from 6 beams, preserve relative paths

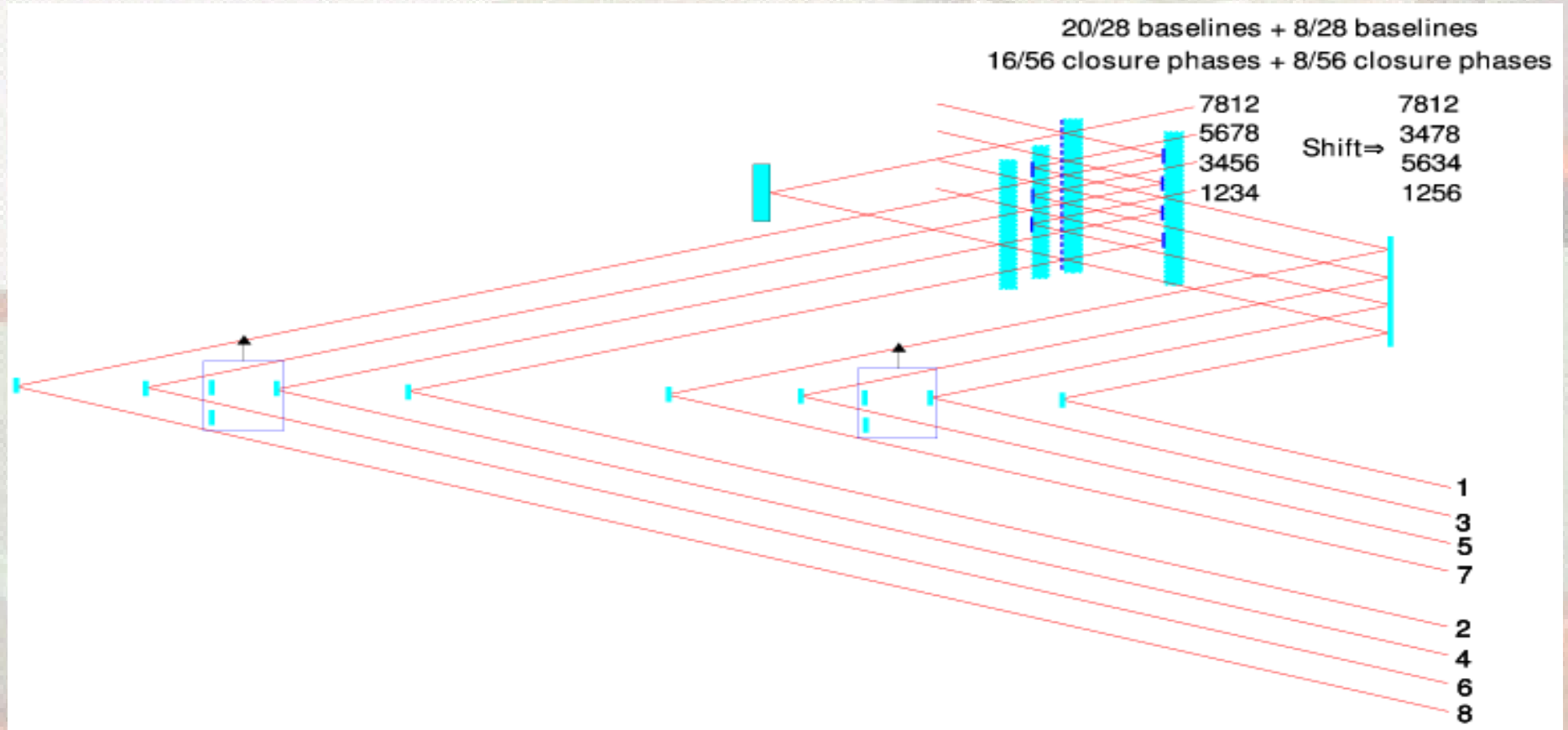


P4(S) Combiner

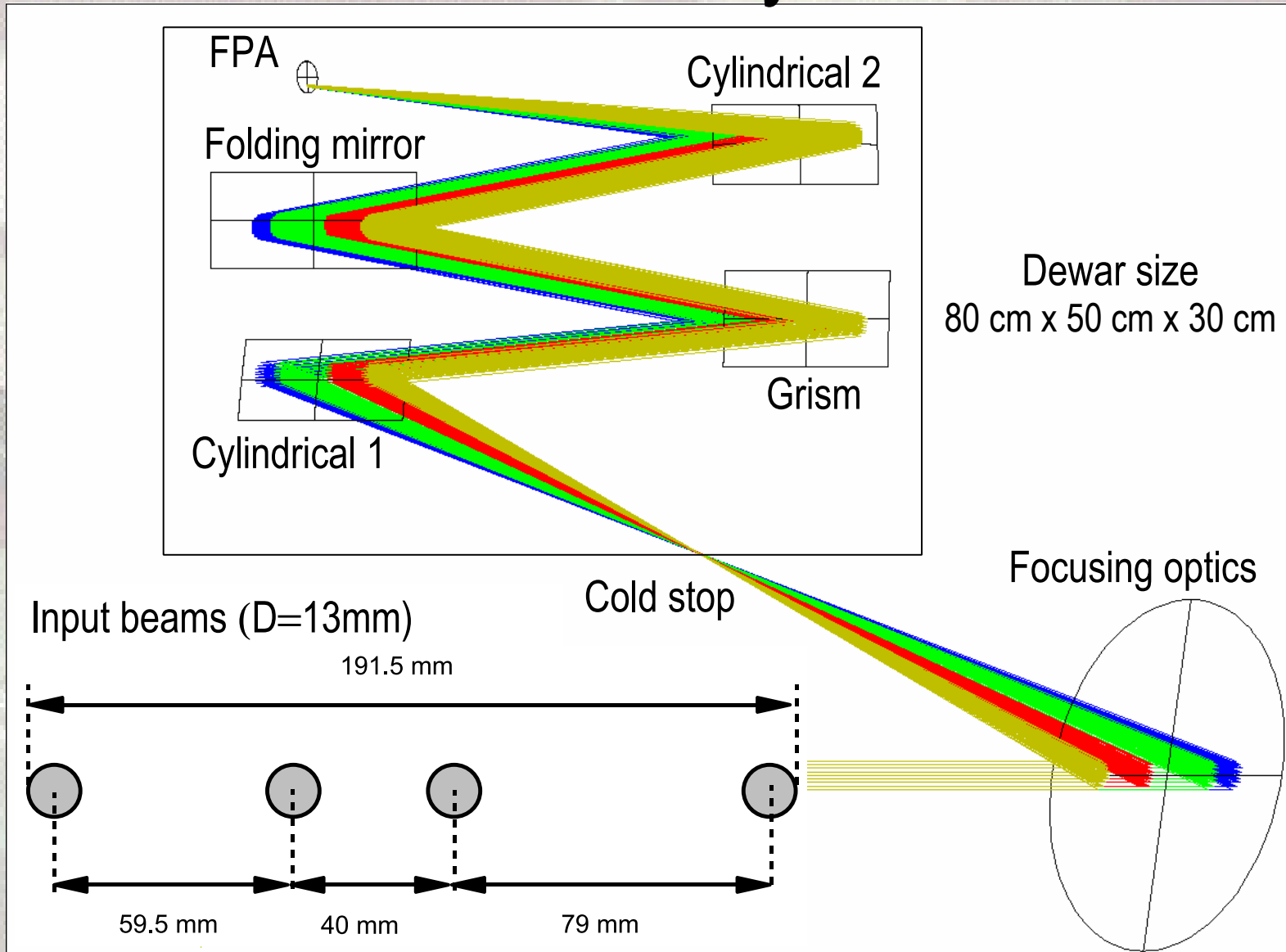
- **Similar to existing contacted 4-way 1-2.5 μm combiner made by Cambridge**
- **Existing combiner features:**
 - Performance verified in lab
 - But designed for 5mm beams
- **Advantages of contacted optics**
 - Stability
 - Ease of alignment
- **Disadvantages**
 - >4 beams unproven until now
 - Need thin or matched coatings



P8 Switchyard and Combiner



I4S beam combiner layout



Spectrographs

- **Spectrograph for pupil plane combiner is simple**
 - Focus single combined beam onto cold stop
 - Inside Dewar : collimate, disperse, and focus spectrum onto FPA
 - Multiplexing possible
- **Spectrograph for image plane is more complex**
 - Same principle but very large anamorphosis factors (>100)
 - Dewar size must be kept reasonable...
- **Reconfigurations**
 - Want to switch:
 - Low \Leftrightarrow Medium dispersion (Science Combiners only)
 - Operating waveband
 - May need change of dispersing element even if not changing spectral resolution
 - OK to switch a few times a night, or even between nights
 - Typically need to adjust last mirror and FPA

Ongoing work

- **Science combiner spectrographs (I... & P...)**
 - Simultaneous optimization for 6 configurations (3 bands \times 2 res.)
- **Pupil plane (P...) combiners**
 - How many combiner outputs can be multiplexed onto each detector ?
 - Contacted combiners – glass procurement
- **Image plane (I...) combiners**
 - Optimization to reduce Dewar sizes
- **Fast switchyards (...S)**
 - Alignment procedure (if needed) after reconfiguration
 - See if repeatable enough to use look-up table

Evaluation of designs

- **Realistic signal-to-noise**
 - FT combiner S/N, sets magnitude limit for operating array
- **Imaging speed, depends on:**
 - Science combiner instantaneous S/N
 - No. of switchyard reconfigurations required (if any)
 - Reconfiguration & calibration overheads
- **Cost : Optics, Mechanics, Detectors**
- **Technical risk**
- **Schedule risk**
- **Calibration and stability issues**

Conclusion

- **Demanding MROI requirements for infrared instruments**
- **Several optical combiners have been designed :**
 - FT : pupil or image plane pairwise, or all-in-one I6 combiner
 - SC : contacted 4/6 beams, or image plane 4/6, switchyards if necessary
- **Most optimizations are finished and final evaluation is ongoing**
- **Milestones:**
 - Evaluation finished in July 06
 - First closure phase in August 08
 - Six telescopes on site in late 2009